

# Latitude-dependent vertical mixing and the tropical thermocline in a global OGCM

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[1] In most ocean general circulation models (OGCM), mixing in the pycnocline is treated with a constant background diffusivity. This creates the following problem. To obtain the observed sharp equatorial thermocline, OGCMs must adopt a pycnocline diffusivity ten times smaller than observed at mid-latitudes. The conflict can only be resolved by switching to a *spatially variable* mixing. In this work we present the GISS mixing model supplemented by Gregg *et al.*'s [2003] finding that the rate of dissipation of internal gravity waves is latitude dependent. We use the new GISS model in a global OGCM and obtain an equatorial thermocline in both the Atlantic and Pacific oceans that is sharper than without the latitude dependence. The model results for the Pacific compare favorably with Kessler's [1990] data. The meridional overturning in the Atlantic and the global poleward heat transport are nearly unchanged from the values obtained without latitude dependence. **INDEX TERMS:** 4231 Oceanography: General: Equatorial oceanography; 4544 Oceanography: Physical: Internal and inertial waves; 4568 Oceanography: Physical: Turbulence, diffusion, and mixing processes; 4522 Oceanography: Physical: El Niño; 4255 Oceanography: General: Numerical modeling. **Citation:** Canuto, V. M., A. Howard, Y. Cheng, and R. L. Miller (2004), Latitude-dependent vertical mixing and the tropical thermocline in a global OGCM, *Geophys. Res. Lett.*, 31, L16305, doi:10.1029/2004GL019891.

## 1. The Problem

[2] A sharp thermocline is both observed [Kessler, 1990] and required to reproduce reasonable ENSO variability [Cane, 1992]. To obtain it, one may adopt a low pycnocline heat diffusivity  $K_h$ . However, if used globally (as in models that employ a constant diffusivity), the resulting meridional overturning is far too weak. Thus, the challenge is to harmonize two contrasting requirements: an *average pycnocline*  $K_h \cong 0.1 \text{ cm}^2 \text{ s}^{-1}$  to obtain a robust meridional overturning (16–25 Sv) [Ganachaud, 2003] and an *equatorial pycnocline*  $K_h$  some 10 times smaller. In this paper, the GISS mixing model, supplemented by the latitude-dependent rate of dissipation of internal gravity waves [Gregg *et al.*, 2003], is shown to satisfy both requirements.

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## 2. The GISS Mixing Model

[3] The GISS mixing scheme [Canuto *et al.*, 2001, 2002] (hereinafter referred to as C1 and C2, respectively) is based on the Reynolds Stress Model and treats temperature and salinity as independent fields, thus incorporating double-diffusion processes. In general, the equations are time dependent and non-local. The solution of the local, stationary model yields the following expressions for the momentum, heat and salinity fluxes:

$$\langle w'u' \rangle = -K_m U_{,z}, \quad \langle w'T' \rangle = -K_h T_{,z}, \quad \langle w's' \rangle = -K_s S_{,z} \quad (1a)$$

We have used the notation  $a_{,z} \equiv \partial a / \partial z$ , brackets indicate a time or ensemble average and a prime denotes a fluctuating field;  $U$ ,  $T$  and  $S$  are the mean velocity, temperature and salinity fields while the  $K_\alpha$ 's are the momentum, heat and salt diffusivities whose form can be given in two representations:

$$K_{\alpha 1} = \Gamma_{\alpha 1} \Lambda^2 \Sigma \quad K_{\alpha 2} = \Gamma_{\alpha 2} \epsilon N^{-2} \quad \Gamma_\alpha = \Gamma_\alpha(Ri, R_\rho) \quad (1b)$$

where  $\Lambda$  is a mixing length,  $\Sigma$  is the mean shear,  $N$  is the Brunt-Vaisala frequency and the  $\Gamma_\alpha$  are mixing coefficients;  $Ri = N^2 / \Sigma^2$  is the Richardson number and  $R_\rho = \alpha S_{,z} / \beta T_{,z}$  is the density ratio; ( $\alpha$ ,  $\beta$ ) are the thermal expansion and haline contraction coefficients. While most models assume a single  $\Gamma = 0.2$ , the GISS model computes all the  $\Gamma_\alpha$  (see Figure 7 of C2). Physically, the first representation in equation (1b) is used when the mixing is due to shear while the second representation is used when the mixing is due to internal wave breaking. In general,  $K_{\alpha 1}$  dominates in the mixed layer while  $K_{\alpha 2}$  dominates in the pycnocline. Since the two mixings may overlap, the physical diffusivity is the sum of the two contributions so as to assure a smooth transition between the two processes:

$$K_\alpha = K_{\alpha 1} + K_{\alpha 2} \quad (1c)$$

The turbulence model provides the  $\Gamma$ 's while an expression for  $\epsilon N^{-2}$  has been provided by several authors [Polzin *et al.*, 1995; Polzin, 1996; Kunze and Sanford, 1996; Gregg *et al.*, 1996; Toole, 1998], and we have adopted the Gregg-Henryey-Polzin parameterization:

$$\epsilon N^{-2} = 0.288 \text{ cm}^2 \text{ s}^{-1} \quad (1d)$$

Even though equation (1d) is a constant,  $K_{\alpha 2}$  is not constant since  $\Gamma_{\alpha 2}$  depends on  $Ri$  and  $R_\rho$ . The GISS mixing model was tested in: 1) 1D ocean models [Burchard and Bolding, 2001], 2) a 3D-OGCM vs. measured heat, salt and

concentration diffusivities (C2); 3) deep convection in the Labrador Sea [Canuto *et al.*, 2004] (hereinafter referred to as C4); 4) inter-comparison with other mixing models in non-convective regimes [Halliwell, 2004], 5) comparison with shear flow data (L. Umlauf, Modeling the effects of horizontal and vertical shear in stratified turbulent flows, submitted to *Deep Sea Research*, 2004).

### 3. Latitude-Dependent Diffusivities

[4] Gregg *et al.* [2003] reported measurements of a latitude-dependent  $\varepsilon(\theta)$  which changes  $K_{\alpha 2}$  to:

$$K_{\alpha 2} = \Gamma_{\alpha 2} \varepsilon_{30} N^{-2} L(\theta, N) \quad (2a)$$

where  $\varepsilon_{30}$  is defined by Gregg *et al.* [2003, equation (4)], while  $L$  is given by:

$$L(\theta, N) = [f \operatorname{Arccosh}(N/f)] [f(30^\circ) \operatorname{Arccosh}(N_0/f(30^\circ))]^{-1} \quad (2b)$$

With the latitude dependence accounted for,  $K_{\alpha 2}$  is now a function of  $Ri$ ,  $R_p$ ,  $N$  and  $f$ .

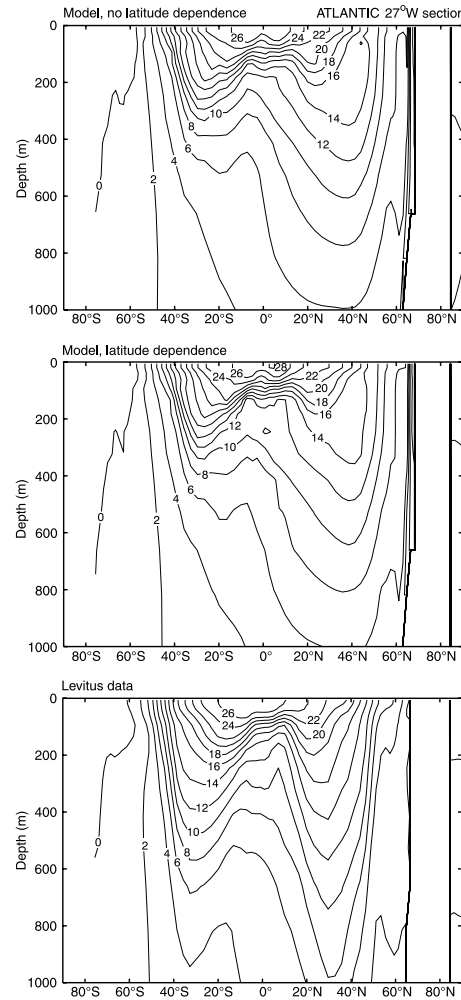
### 4. Global Tests

[5] We used the 3D NCAR-CSM ocean model with  $3^\circ \times 3^\circ$  horizontal resolution, 25 vertical levels (7 levels in the upper 200 m) and a maximum depth of 5 km. We used the GM [Gent and McWilliams, 1990] parameterization for mesoscales with eddy advection and isopycnal diffusion coefficients equal to  $8 \bullet 10^2 \text{ m}^2 \text{ s}^{-1}$ . The OGCM was integrated to equilibrium for 126 years using split tracer and momentum time steps, with tracer time step ten times momentum throughout (see C1 and C4). Since Gregg *et al.* found diffusivities larger than the molecular value at the equator, we imposed a minimum of 0.07 on  $L$  in equation (2b). The global annually averaged temperature and salinity profiles with latitude dependence show little change compared to Figures 5–6 of C4 with no latitude dependence. The meridional overturning and poleward heat transport are hardly affected; the Atlantic meridional stream function is 1% larger than without latitude-dependent mixing and the northward heat transport for the entire ocean is reduced from 1.36 Pw to 1.25 Pw which are within the range of uncertainties of the data [Macdonald and Wunsch, 1996].

### 5. Tropical Atlantic

[6] In Figure 1 we show the Atlantic thermocline at  $27^\circ \text{W}$ , a location in the middle of the Atlantic away from land masses. We found similar results at  $20^\circ \text{W}$  close to Africa. Figure 1 shows model results together with Levitus data [Antonov *et al.*, 1998]. A sharpening of the thermocline in the middle panel is visible. To illustrate the difference, we computed the degree of stratification due to temperature,  $N_h^2 = N^2(1 - R_p)^{-1}$ . At the Equator and  $27^\circ \text{W}$ , at 62.5 m, the GISS model (no latitude dependence) yields  $N_h^2 = 2.4 \cdot 10^{-4} \text{ s}^{-2}$  and  $N_h^2 = 2.8 \cdot 10^{-4} \text{ s}^{-2}$  (with latitude dependence), indicating a sharpening. At  $20^\circ \text{W}$ , the corresponding values are  $N_h^2 = 2.2 \cdot 10^{-4} \text{ s}^{-2}$  and  $2.6 \cdot 10^{-4} \text{ s}^{-2}$ , showing the same trend.

[7] In Figure 2, we show the annually averaged heat diffusivity in the Atlantic thermocline at  $27^\circ \text{W}$  from  $45^\circ \text{S}$  –



**Figure 1.** Annually averaged temperature sections at  $27^\circ \text{W}$  in the Atlantic with and without the latitude dependence. The two upper panels correspond to the GISS mixing model without and with latitude dependence while the lowest panel shows the Levitus data [Antonov *et al.*, 1998].

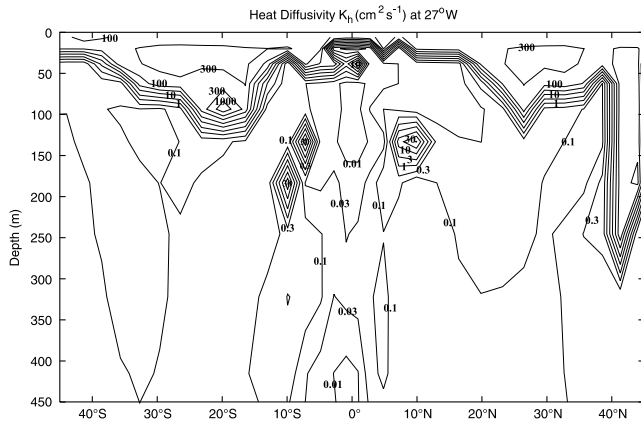
to  $45^\circ \text{N}$ . Diffusivity varies widely with both latitude and depth. Values larger than  $30 \text{ cm}^2 \text{ s}^{-1}$  are due to weakly unstable gradients  $N^2 \sim -10^{-6} \text{ s}^{-2}$  and not internal wave breaking. Similar instabilities occur in the case without latitude dependence. The instabilities and the large associated diffusivities are not present at all times but persist for several days at a time and recur frequently so they contribute to the annual average. Diffusivities  $> 10^2 \text{ cm}^2 \text{ s}^{-1}$  are typical of the mixed layer [Caldwell *et al.*, 1997]. A z-variation of measured diffusivities [Peters *et al.*, 1988] shows a range almost as large as the one presented here.

### 6. Tropical Pacific

[8] In most mixing models  $K_{\alpha 1}$  is assumed to be a function of only the Richardson number  $Ri$  while  $K_{\alpha 2}$  is taken to be a constant *background diffusivity*:

$$K_{\alpha 1} = f_{\alpha}(Ri) \quad K_{\alpha 2} = \text{constant} \quad (3a)$$

The values of  $K_{\alpha 2}$  for momentum and heat are  $K_{m2} = 1$ ,  $K_{h2} = 0.1$  [Pacanowski and Philander, 1981] (PP);  $K_{m2} =$



**Figure 2.** Annually averaged, latitude dependent heat diffusivity  $K_h$  ( $\text{cm}^2\text{s}^{-1}$ ) at  $27^\circ\text{W}$  in the Atlantic.

0.2,  $K_{h2} = 0.01$  [Peters *et al.*, 1988] (PGT);  $K_{m2} = 1 - 2$ ,  $K_{h2} = 0.5 - 1$  [Gent, 1991];  $K_{m2} = 0.01$ ,  $K_{h2} = 0$  [Latif *et al.*, 1994; Frey *et al.*, 1997] (MPI);  $K_{m2} = 1$ ,  $K_{h2} = 0.1$  [Chen *et al.*, 1994] (Hybrid Vertical Mixing (HVM));  $K_{m2} = 1$ ,  $K_{h2} = 0.1$  [Large and Gent, 1999] (KPP);  $K_{m2} = 2$ ,  $K_{h2} = 0.2$  [Yu and Schopf, 1997]; and  $K_{m2} = 0.2$ ,  $K_{h2} = 0.01$  [Wilson, 2000, 2002] (Integrated Power PGT model (IP)). The functions  $f_\alpha(\text{Ri})$  can be found in the references cited.

[9] Wilson [2000, 2002] analyzed four of the eight schemes, namely PP, PGT, MPI and IP and concluded that: 1) only PP is unable to yield a sharp thermocline and 2) that a much reduced set of values  $K_{m2} = 0.0134$ ,  $K_{h2} = 0.1$   $K_{m2}$  (called the MPP, modified PP scheme) is needed to yield a sharp thermocline. The KPP model was tested by Large and Gent [1999] at four locations  $165^\circ\text{E}$ ,  $170^\circ\text{W}$ ,  $140^\circ\text{W}$  and  $110^\circ\text{W}$ . In the first two locations, the model reproduced the temperature profiles from TOGA-TAO while in the last two (corresponding to ENSO regions), the predicted thermocline was too diffuse, a behavior that Large and Gent attributed not to the vertical mixing model but to the uncertainties in

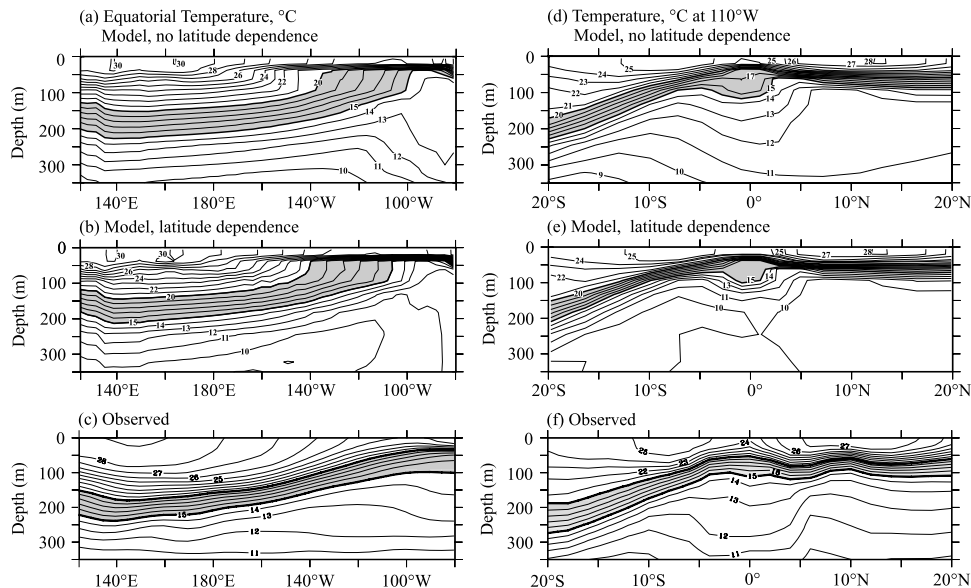
the wind stress and heat flux forcing. The YS model at  $110^\circ\text{W}$  also yields a diffuse thermocline not very different from that of KPP. In the HVM model, the authors test three diffusivity models: a) a constant 50 m mixed layer plus Gent's [1991] Ri-dependent model, b) a bulk mixing scheme and c) their own HVM. At  $110^\circ\text{W}$ , the HVM yields a slightly sharper thermocline than the other two models but still not in full accordance with the data. In summary, the above “background” values fall in the interval:

$$0.01 \leq K_{h2} (\text{cm}^2\text{s}^{-1}) \leq 1 \quad (3b)$$

The lower value can yield a sharp thermocline but it is larger than the smallest value of Gregg *et al.* [2003] and if used globally, it yields a very weak meridional overturning. In Figure 3 we compare the longitudinal and latitudinal sections of the equatorial Pacific thermocline temperatures from the GISS model with the measured data [Kessler, 1990]. Along the Equator (left panels) the sharpening of the thermocline between  $15-20^\circ\text{C}$  is small, while the meridional sections show a somewhat greater improvement over the no-latitude dependence case. To quantify this behavior, in the zonal section, the distance between the  $15-20^\circ\text{C}$  isotherms at  $160^\circ\text{W}$  is 69 m (no latitude dependence) but 65 m (latitude dependence). In the meridional section, the distance between the same isotherms at the equator is 83 m (no latitude dependence) and 68 m with latitude dependence. The bulging out of isotherms has been reduced in the latitude-dependent case.

## 7. Conclusions

[10] In most mixing models, the pycnocline diffusivity is treated as a globally uniform, background value. The vertical diffusivity required for realistic meridional overturning and poleward heat transport [Bryan, 1991] results in an unrealistically diffuse tropical thermocline that inhibits ENSO variability [Lau *et al.*, 1992; Miller and Jiang, 1996]. Because of the global influence of ENSO [Yulaeva and



**Figure 3.** Kessler [1990] equatorial Pacific data vs. model predictions.



Wallace, 1994], this would distort the circulation of a coupled AGCM worldwide. The GISS model with latitudinal dependence of internal gravity wave dissipation [Garrett, 2003; Gregg *et al.*, 2003] yields both a robust North Atlantic stream function and a sharp thermocline. It may be of interest to notice that the latter was obtained using only 7 levels in the upper 200 m, a modest number compared to typical ENSO simulations with an OGCM [Philander *et al.*, 1992]. The present results are an attempt to account for a spatially variable turbulent kinetic energy dissipation rate  $\epsilon$  in an OGCM. We have accounted for the latitude dependence in the pycnocline but an additional source of spatial variability represents tidal energy dissipation in the ocean's bottom [Simmons *et al.*, 2004]. Once the latter is added to the GISS model, we hope to obtain an even more complete representation of mixing from the surface to the ocean's bottom. Work in that direction is under way.

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